stellar interferometry : an overview about basics

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stellar interferometry : an overview about basics

sections

- introduction : a problem raised
- science context and motivation
- few academic reminders
- basics for interferometry and aperture synthesis
- Iimitations and subsequent needs
- interferometers : principle, production, typology
- difficulties in real world (and some remedies)
- managing with data and some results
- quick-look at some alternative HAR methods
- nulling interferometry and coronagraphy

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just talking (where are we ??)

time-signals can be described

either by amplitudes (signal) or by frequencies (spectrum)

similarly

there are two ways to describe brightness distributions : the direct description, based on coordinates (image) the other one, based on spatial frequencies (spatial spectrum)

For brightness distributions, when images are not at hand a "spectrum analyser" must be found Once available, coming back from the spectrum to the image could be possible.

This is the goal of aperture synthesis

Thus we have now to conceive and built such a spectrum analyser

an important auxiliary : Fourier optics

- key protagonists : complex amplitude and wavefront
- building the tool :
 - a very quick look
 - Huyghens Fresnel principle
 - here comes Fourier formalism
- ☐ illustrative example(s)

complex amplitude of fields

source S, observation point P

Wave at S
$$V_{s}(t) = A.exp(i. 2\pi.v.t)$$

At P, same behaviour but time-delayed $V_P(t,x,y,z) = A.exp(i.2\pi.v.(t-r/c))$



for imaging, relevant information is in the phase distribution $2\pi .r/\lambda$ at each observation point, so the useful description is given by the complex amplitude

It describes the shape of the wavefront (equiphase surface)

 $V_{P}(t,x,y,z) = A.exp(i.(v. t - r/\lambda))$





the phase mathematically describes the shape of the wavefront wavefronts are "equiphase surfaces"





summary : if T know w(

if I know $\psi(\xi,\eta,0)$ amplitude transmitted by "s" (pupil plane) I can calculate $\psi(x,y,z)$ amplitude over screen O (image plane) Intensity at O = squared modulus of $\psi(x,y)$

tools to do that : approximations + Huyghens-Fresnel principle

Huyghens-Fresnel principle

complex amplitude

already seen : as the field propagates the phase increases according to length of traveled path

Huyghens Fresnel principle:

every point Q_n within an amplitude distribution emits a spherical wave, all waves are synchronous but not necessarily "in phase"

The field seen by P is the sum of the amplitudes of the spherical waves (complex numbers)

$$\psi(P) = \sum \psi(Q_n) . exp(i.\frac{2\pi}{\lambda}.path))$$

$$\psi_z(x,y) = \int \psi_0(\xi,\eta) . exp(i.\frac{2\pi}{\lambda}.r(\xi,\eta,x,y)) . d\xi.d\eta$$















other useful auxiliaries :

linear filtering and transfer function



bi-dimensional situation

object-image relationship in the direct space (coordinates space)

I(x,y) = O(x,y) * h(x,y)
h(x,y) = response to dirac : Point Spread Function (typical : Airy pattern)



description in the frequencies space (Fourier space)

$$\hat{I}(u,v) = \hat{O}(u,v).\hat{h}(u,v) = \hat{O}(u,v).T(u,v)$$

T(u,v) = transfert function = FT of impulse response shows how frequencies are transmitted : gain (complex) Also : T(u,v) shaped like autocorrelation of pupilla (Rayleigh theorem)



 D/λ : cut-off frequency. Consequence : higher frequencies are lost





fundamental principles :

Coherence and the second key : VCZ theorem



defining coherence : a mutual notion

coherence of fields is the ability they might have to produce observable interferences when mixed (duration equal or larger than the detector integration time say 10^{-6} s)

coming back to our first tool

quadratic detection, energy detection (visible and infrared)

incoming field Ψ detection process: output = $\langle |\psi|^2 \rangle_{\tau}$ τ = integration time,

notation <> τ means "averaging over τ "



"mixing fields" means ψ_1 and ψ_2 arriving together on the detector $<|\psi_1 + \psi_2|^2 > = <|\psi_1|^2 > + <|\psi_2|^2 > + 2.\operatorname{Re}(<\psi_1 \cdot \psi_2 >)$ $<|\psi_1 + \psi_2|^2 > = I_1 + I_2 + 2.\operatorname{Re}(<\psi_1 \cdot \psi_2 >)$ energy terms interference term the single cosine model is not compatible with observation of interferences



need another model for wave of light, a model with random features

a relevant model :

train of damped oscillations with random emission time and phase at origin

length of a wagon : $\tau_{\rm c} \approx$ 1/ spectral $\Delta \nu$

how to quantify this coherence ability? _2

$$<|\psi_1 + \psi_2|^2 > = I_1 + I_2 + 2.\operatorname{Re}(<\psi_1, \psi_2>)$$

< $\psi_1.\psi_2^*$ > is something like a covariance it could serve but depends on incoming amplitudes (leading to intensities) better to define

a dimensionless quantity free from differences in intensity

A complex number which modulus varies between 0 and 1 is appropriate

complex degree of coherence
$$\gamma_{12} = \frac{\langle \psi_1 \cdot \psi_2^* \rangle}{\sqrt{I_1 \cdot I_2}}$$
 with $I_k = \langle |\psi_k|^2 \rangle$

 γ_{12} can be seen as a complex normalized $\mbox{ covariance }$

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about addition of fields : generic situations _ 1
incoherent addition
the interference term is destroyed
by averaging over the random phases

$$I = \langle |\psi_1 + \psi_2|^2 \rangle = \langle \left| \sum_k \psi_{1k} + \sum_n \psi_{2n} \right|^2 \rangle S_2 \xrightarrow{M_{\text{thereform}}} V_2$$

$$I = \langle \left| \sum_k \psi_{1k} \right|^2 \rangle + \langle \left| \sum_n \psi_{2n} \right|^2 \rangle + 2.Re\left[\langle \sum_k \sum_n \psi_{1k} \cdot \psi_{2n}^* \rangle \right] \rangle$$
all $\langle \psi_{1k} \cdot \psi_{2n}^* \rangle = convey \langle exp[i(\phi_k - \phi_n)] \rangle$ and averaging results in zero

 $I = I_1 + I_2 + nothing !$

two distinct sources cannot produce an observable interference term

consequence : introducing a non-coherence relationship

$$<\psi(\alpha).\psi^{\star}(\beta)>=<|\psi(\alpha)|^{2}>.\delta(\alpha-\beta)$$

$$S_{1} \qquad \psi_{1}$$

$$S_{2} \qquad \beta \qquad \psi_{2}$$

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about addition of fields : generic situations _ 2

coherent addition

let us consider a machine performing addition of field collected at P1 and P2 from a unique point-like source and nearly equal paths r_1 and r_2



each "wagon" is like :
$$A(t - t_k) exp\left[i \cdot \frac{2\pi}{\lambda}r_{1 or 2} + i \cdot \phi_k\right]$$

and interference term conveys components like :

<
$$A(t - t_k - \frac{r_1}{c}).A(t - t_k - \frac{r_2}{c}).exp\left[i.\frac{2\pi}{\lambda}.(r_1 - r_2) + i(\phi_k - \phi_k)\right]$$
 >

now averaging over ϕ has no effect, and interference term remains, in spite of the effect of statistics. zero

However it is non-zero only if "wagons" are overlaping what requires that (r1-r2) must be smaller than the length of a wagon (subsequently named : coherence length)

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path difference must be small enough



a third situation : synthesis of the two previous ones _1

addition partially coherent several sources $S_1, S_2, ..., S_n$ and the addition machine



phenomenology

all sources Sk yield an observable interférence term (situation 2), independantly one to another, each term depends on $(r_{k1} - r_{k2})$ in other words depends on its location



the resulting value for the interference term depends on the (angular) extension of the distribution of point-like sources a third situation : synthesis of the two previous ones_2

On P₁ and P₂ comes the sum (over α) of fields from directions (α) respective paths are r₁(α) and r₂(α)

$$\psi_1 = \int \psi(\alpha) . exp(-i . \frac{2\pi}{\lambda} . r_1(\alpha)) . d\alpha \quad \psi_2 = \int \psi(\alpha) . exp(-i . \frac{2\pi}{\lambda} . r_2(\alpha)) . d\alpha$$

then we have (classical expression for a product of integrals)

$$<\psi_{1}.\psi_{2}^{*}>=\int\int\left(<\psi(\alpha).\psi^{*}(\beta)>\exp(-i.\frac{2\pi}{\lambda}.[r_{1}(\alpha)-r_{2}(\beta)]).d\alpha.d\beta\right)$$

look !

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So what?

$$\langle \psi_1.\psi_2^* \rangle = \int \int \langle \psi(\alpha).\psi^*(\beta) \rangle exp(-i.\frac{2\pi}{\lambda}.[r_1(\alpha) - r_2(\beta)]).d\alpha.d\beta$$

and thanks to non-coherence relation
 $\langle \psi_1.\psi_2^* \rangle = \int \int O(\alpha).\delta(\alpha - \beta).exp(-i.\frac{2\pi}{\lambda}.[r_1(\alpha) - r_2(\beta)]).d\alpha.d\beta$
and thanks to dirac properties
 $\langle \psi_1.\psi_2^* \rangle = \int O(\alpha).exp(-i.\frac{2\pi}{\lambda}.[r_1(\alpha) - r_2(\alpha)]).d\alpha \int \delta(\alpha - \beta).d\beta$
and so and so?
 $\langle \psi_1.\psi_2^* \rangle = \int O(\alpha).exp(-i.\frac{2\pi}{\lambda}.[r_1(\alpha) - r_2(\alpha)]).d\alpha \int \delta(\alpha - \beta).d\beta$
isn't that beautiful? Fourier strikes back
 $\langle \psi_1.\psi_2^* \rangle = \int O(\alpha).exp(-i.\frac{2\pi}{\lambda}.[B.\alpha]).d\alpha$

last step towards complete happiness

a by-product: $<\psi_1.\psi_1^* > = < |\psi_1|^2 > = I_1 = \int O(\alpha).d\alpha = \hat{O}(u=0)$ same for ψ_2

let us then recast the degree of coherence

$$\gamma_{12} = \frac{\langle \psi_1 \cdot \psi_2^{\star} \rangle}{\sqrt{I_1 \cdot I_2}} = \frac{\hat{O}(\frac{B}{\lambda})}{\hat{O}(0)}$$

in other words with $P_1P_2 = B$ and with λ the degree of coherence is the FT at frequency B / λ of the angular brightness distribution of the source normalized on its FT at origin

this simply is the Van Cittert & Zernike theorem





soufflons un peu, et prenons l'air avant la suite qui est pire que tout !

Long Baseline Interferometry

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functional features

interferometry : measuring by means of interferences

among several compact descriptions the two main features here to handle are :

- interferometer = machine performing coherent addition
- interferometer = filter for spatial frequencies

but the main feature for science concerns is : (simply a consequence of coherent addition)

interferometer = instrument making observable the degree of coherence

here we look at the academic case (Fizeau configuration)

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interferometer makes observable the degree of coherence

generally with two apertures, only V can be extracted from data ϕ_{12} is mixed with random spurious phases and is not available

So the observation here, only yields V = modulus of γ_{12} also called : amplitude of fringes

note : with point-like source V = 1_____

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interferometer as filter for spatial frequencies

point-like source yields impulse response

$$I(x) = 2.I_0.Airy.\left[1 + \cos\left(2.\pi \cdot \frac{B}{\lambda} \cdot \frac{x}{F}\right)\right] = 2.I_0.h(x)$$

transfer function (FT of h(x) normalized to unity)



any source yields: I(x) = O(x) * h(x) and $\hat{I}(u) = \hat{O}(u).T(u)$

interferometer allows sensing the source spectrum at frequencies as high as B/λ

note : two telescopes, one baseline, one spatial frequency

back to our "fringed lobe" : what is it doing?

the lobe questions the source about the presence of a particular spatial frequency (the one born by the fringes : B/ λ)

it makes a measure of this "presence",

in extracting info from the spatial spectrum

conceptually the source answer is the so-named "visibility" of the source at frequency B/λ

The visibility is given par by the modulation rate of the observed fringes on the camera

thanks to VanCittert & Zernike, we now know that this visibility is the degree of coherence of the source at frequency B/λ

u1 📘

note : one baseline, one frequency, each time a little piece of information (a component of the spatial spectrum)



u1, how much ?





As announced, we have picked up information in the frequency space